

INTERACTION of PARTICLES and RADIATION with MATTER

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- INTRODUCTION
- PLAYGROUND
- HOW TO DETECT ?
- RADIATION
- MASSIVE PARTICLES

INTRODUCTION

PLAYERS in the SUB-ATOMIC WORLD

PLAYERS I: PARTICLES

particle



detector



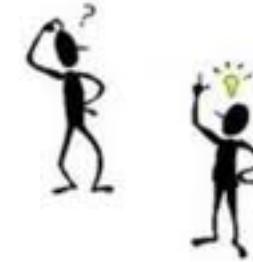
registration



Light



Heavy



PARTICLES

What characterizes a particle?

mass

m *includes $m = 0$*

charge

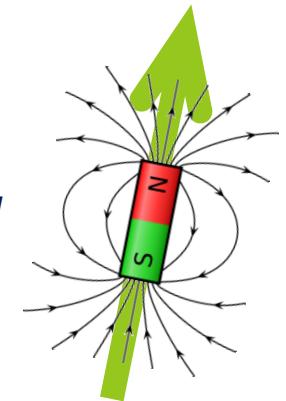
Q

Spin *intrinsic angular momentum*

S \Leftrightarrow *magnetic moment μ*

life time

τ_0



size *for non elementary particles*

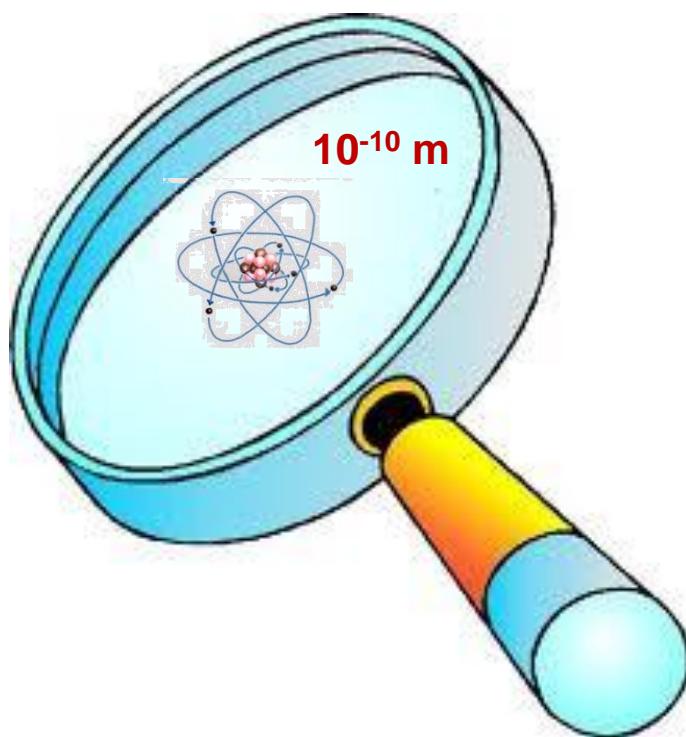
$\sqrt{\langle r^2 \rangle}$ *root-mean square radius*

quantum mechanics

$\lambda = \frac{h}{p}$ *matter waves (de Broglie 1922)*

PARTICLES in 1932

atoms



	atomic shells	nucleus	
	electron e	proton p	neutron n
Q	- 1	+ 1	0
M	$M_p / 1836$	M_p	$\approx M_p$
S	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$
size	$< 10^{-18} \text{ m}$	$0.8 \cdot 10^{-15} \text{ m}$	
life time τ_0	$> 10^{26} \text{ y}$	$> 10^{29} \text{ y}$	886 s
decay	-	-	$n \rightarrow p e^- \bar{\nu}$

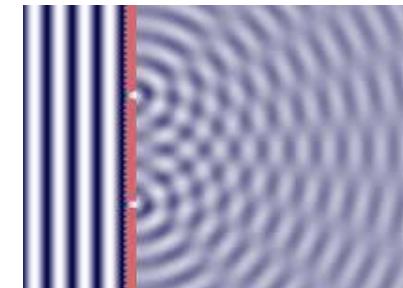
PARTICLES ...

new particles – unstable being free

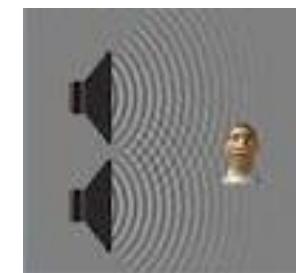
	pions	kaons	<i>many more</i>
	π	K	...
Q	$0, \pm 1$	$2 \times 0, \pm 1$	
M	$\approx M_p / 7$	$\approx M_p / 2$	
S	0	0	$0, \frac{1}{2}, 1, \frac{3}{2}, 2, \dots$
size	$0.6 \cdot 10^{-15} \text{ m}$	$0.6 \cdot 10^{-15} \text{ m}$	
life time τ_0	π^\pm $26 \cdot 10^{-9} \text{ s}$ π^0 $8 \cdot 10^{-17} \text{ s}$	K^\pm $12 \cdot 10^{-9} \text{ s}$ $K^0_{S,L}$ $9 \cdot 10^{-10} / 5 \cdot 10^{-8} \text{ s}$	
decay	$\pi^\pm \rightarrow \mu^\pm \nu$ $\pi^0 \rightarrow \gamma \gamma$	$K^\pm \rightarrow \mu^\pm \nu, \dots$ $K^0 \rightarrow \pi^+ \pi^-, \pi^0 \pi^0, \dots$	

PLAYERS II: WAVES

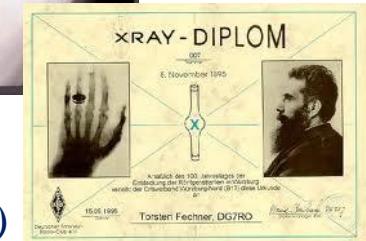
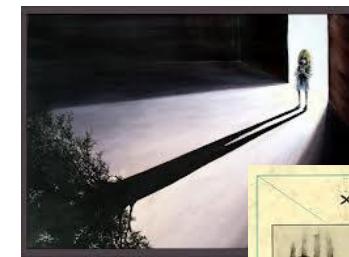
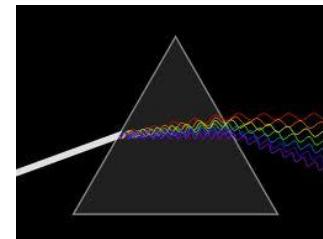
fluid



gas



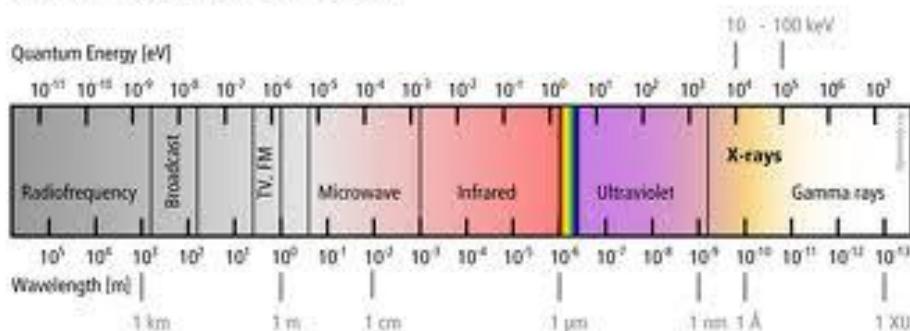
„light“



fundamental constant: $c = \text{speed of light in vacuum} (\cong 30 \text{ cm / ns})$

ELECTROMAGNETIC RADIATION

The Electromagnetic Spectrum



wave length

λ

frequency

ν

wave propagation velocity in vacuum

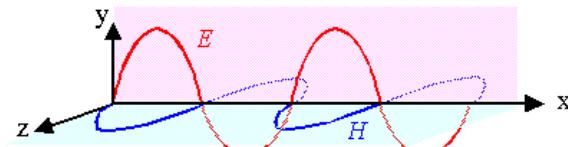
“ “ “ in medium

index of refraction

$$c = \lambda \nu$$

$$c' = \lambda' \nu < c$$

$$n = c / c'$$

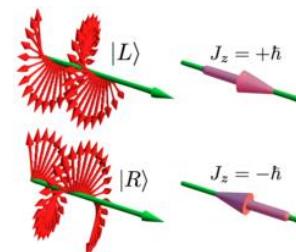


quantum mechanics: waves can be particles

Photon

$S = 1$

$m = \pm 1, \text{ no } m = 0$



*left/right
circular
polarization*

having energy

$$E = h\nu = \frac{hc}{\lambda}$$

(Einstein 1905)

COMPARISON

massive particles

total energy

$$\begin{aligned} E_{\text{total}} &= \sqrt{p^2 c^2 + m_0^2 c^4} \\ &= \gamma m_0 c^2 \\ T_{\text{kin}} &= E_{\text{total}} - m_0 c^2 \end{aligned}$$

rest mass

$$m_0 \neq 0 \quad \textit{range in matter}$$

charge

$$Q \neq 0 \quad \textit{deflection in el.-mag fields}$$

life time

$$\tau = \gamma \tau_0 \quad \textit{decay length } l = v \tau$$

el.-mag. radiation

$$E_{\text{total}} = pc$$

$$= h\nu \quad h \textit{ Planck constant}$$

$$= \hbar \omega \quad h = \textit{minimal action} \quad \left(\hbar = \frac{h}{2\pi} \right)$$

$$= 0 \quad \textit{attenuation in matter}$$

$$= 0 \quad \textit{no deflection}$$

$$= \infty$$

relativistic factor

$$\gamma = \frac{1}{\sqrt{1-\beta^2}}, \quad \beta = \frac{v}{c}$$

$$\lim_{v \rightarrow c} \gamma \xrightarrow{} \infty$$

PLAYGROUND

by means of „normal“ matter

ATOMS and NUCLEI I

Each substance is composed of chemical elements ≡ basic set of atoms (Dalton, ...)



1 Mol contains always the same number of particles $N_A = 6 \cdot 10^{23}$

The ratios of molar masses of the elements are almost ratios of integer numbers.

Atomic mass unit (a.m.u.)

$$1 \text{ a.m.u.} = \frac{\mathbf{m}(\text{C atom})}{12} = \frac{12 \text{ g}}{12 \cdot N_A} = 1.66 \cdot 10^{-27} \text{ kg}$$

Periodic system of elements

Mendeleev: Ordering scheme according to chemical properties

ATOMS and NUCLEI II

Discovery of the electron (J.J. Thomson 1897)

Interpretation of the rays accompanying gas discharges:

Kathode rays: Electrons (negatively charged and light)

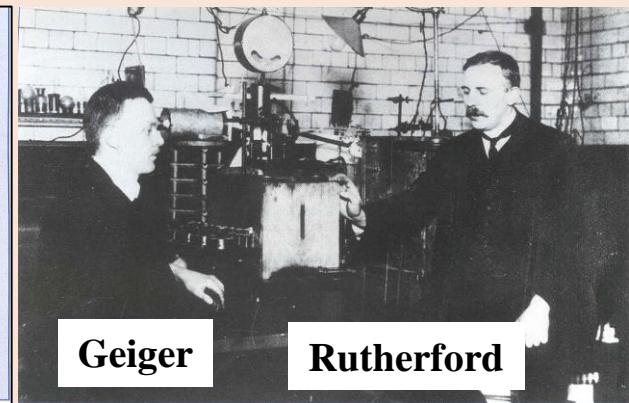
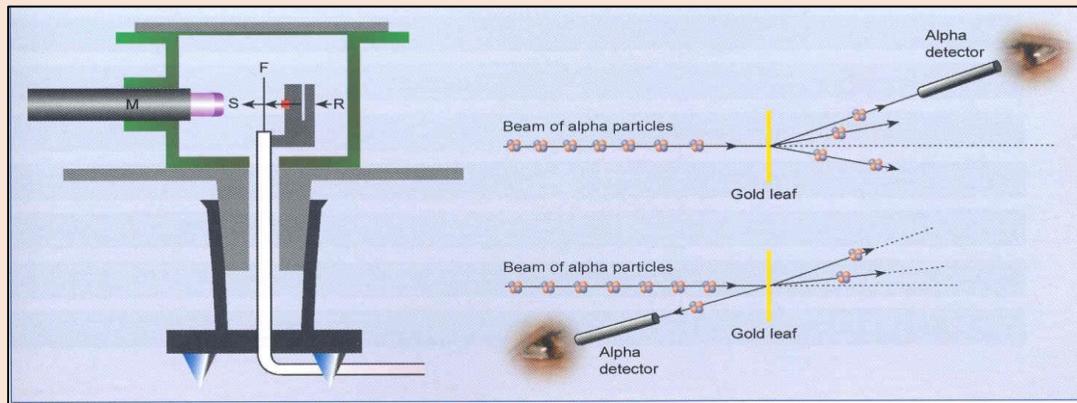
Channel rays: Ions (positively charged and heavy)

$$\mathbf{m}_e = \frac{1}{1836} \cdot \mathbf{m}_u, \quad q_e = -e$$

Discovery of the atomic nucleus (Geiger & Rutherford 1911)

Almost the full mass of an atom of diameter $\sim 10^{-10} \text{ m}$ is concentrated in a tiny volume of radius $\sim 10^{-15} \text{ m}$

Proof: Collision kinematics – backscattering only from heavy collision partners



Geiger

Rutherford

\Rightarrow total atom is electrical neutral: Atomic nucleus - electric charge $+ Z q_e$ Z positive integer
Atomic shell - electric charge $- Z q_e$

ATOMS and NUCLEI III

Bohr-Sommerfeld model of atoms

main quantum number $n = 1, 2, \dots$

“*main shell*”

angular momentum $\ell = 0, \dots, n-1$

„*sub-shell*“

magnetic quantum number $m = -\ell, -\ell+1, \dots, \ell-1, \ell$

($2\ell + 1$) possible orientations of angular momentum vector

in external field

intrinsic electron spin $S = \frac{1}{2}$

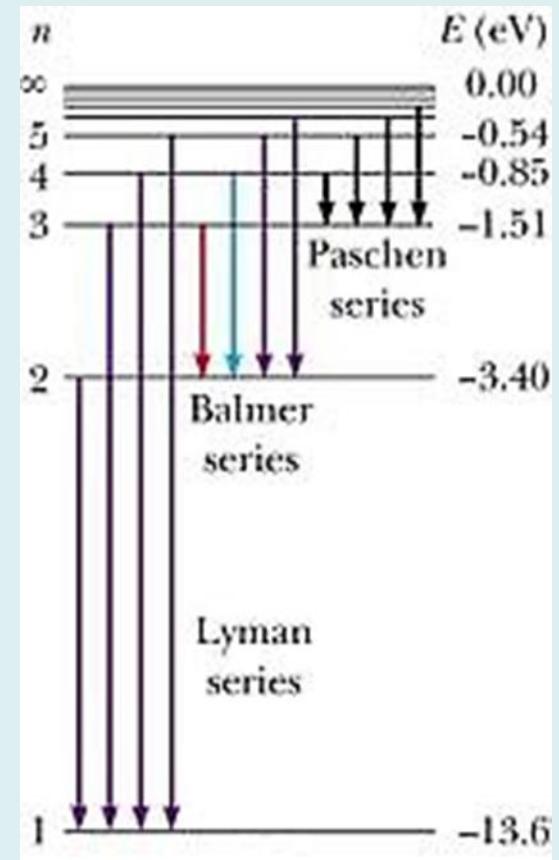
intrinsic angular momentum

2 possible orientations of spin vector *in external field* $S = \pm \frac{1}{2}$

total spin: $\vec{j} = \vec{\ell} + \vec{S}$ $|\vec{j}| = \frac{1}{2}, 1, \frac{3}{2}, \dots$

ℓ and S are measured in units \hbar

main shells for Z = 1 (H)



ATOMS and NUCLEI IV

$S = \frac{1}{2}$ particles are called *fermions*

Pauli principle:

Only one fermion is allowed in a particular quantum state: For atoms = (n, ℓ, m, S)

\Rightarrow maximum no. of electrons per sub-shell: $2 \cdot (2\ell + 1)$

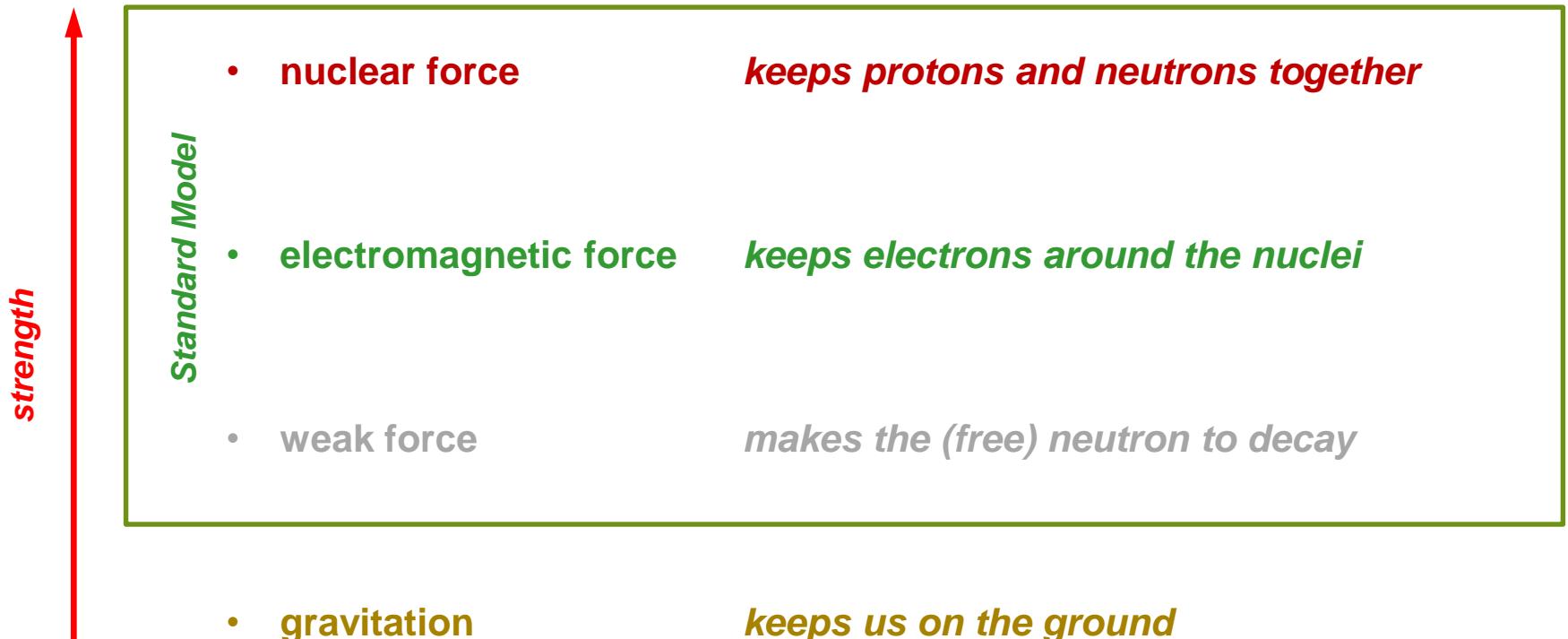
Periodic system of elements

Mendelejev: Ordering scheme according to chemical properties

$$A(Z,N) + Ze^-$$

*outmost
incomplete
shell
determines
chemistry*

FORCES



HOW TO DETECT ?

,,via“ electric force !

SIGNAL CREATION



- **access via electric charges**
- **produced in elementary processes with atoms and nuclei**

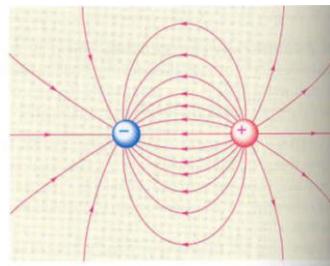
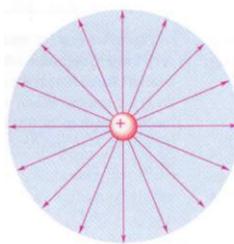
EL.-MAG. INTERACTION I - ELECTRIC FORCE

the force is mediated in the

classical picture

by field around a source

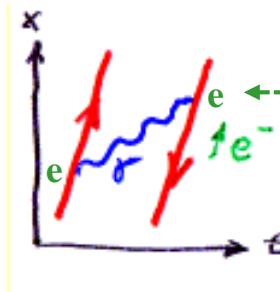
$$\mathbf{F}_{\text{Coulomb}} = \frac{1}{4\pi\epsilon_0} \cdot \frac{\mathbf{Q}_1\mathbf{Q}_2}{r^2}$$



quantum world

field quanta = „virtual“ particles

„light“ particles = photons γ



$$\begin{aligned}\text{coupling } \alpha &= \frac{e^2}{4\pi\epsilon_0\hbar c} \\ &= \frac{1}{137}\end{aligned}$$

electromagnetic radiation = E and B fields interact with electric charges

EL.-MAG. INTERACTION II - CHARGES in EL.-MAG. FIELDS

- electric field

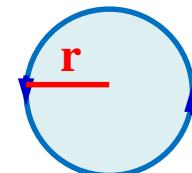
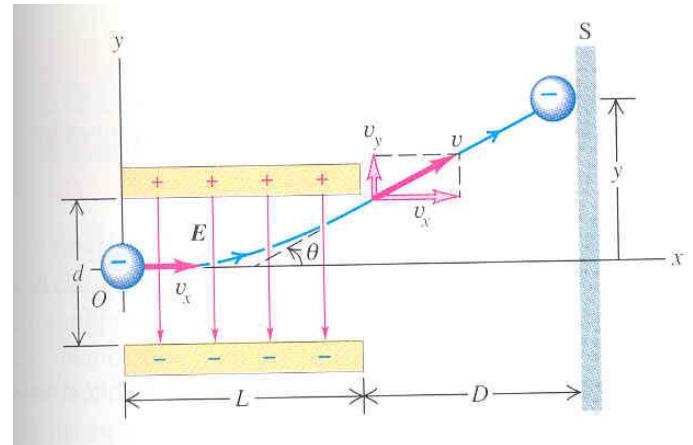
$$\vec{F} = m\ddot{\vec{x}} = Q \cdot \vec{E}$$

$$\Rightarrow \frac{Q}{M}$$

- magnetic field

$$\vec{F} = m\ddot{\vec{x}} = Q \cdot (\vec{v} \times \vec{B})$$

$$\Rightarrow p$$



$B = \text{const.}$
 \Rightarrow *circular motion*
 $B \perp$ *plane of projection*

$$\omega = \frac{Q}{M} B \qquad \omega = \frac{2\pi}{T}$$

$$mv^2 / r = Q \cdot v \cdot B$$

$$p = Q \cdot B \cdot r$$

INTERACTION OF

ELECTRO – MAGNETIC RADIATION

SCATTERING

EXCITATION - SCINTILLATORS produce “LIGHT”

PHOTO EFFECT

COMPTON EFFECT

PAIR PRODUCTION

BREMSSTRAHLUNG & EL.-MAG. SHOWER

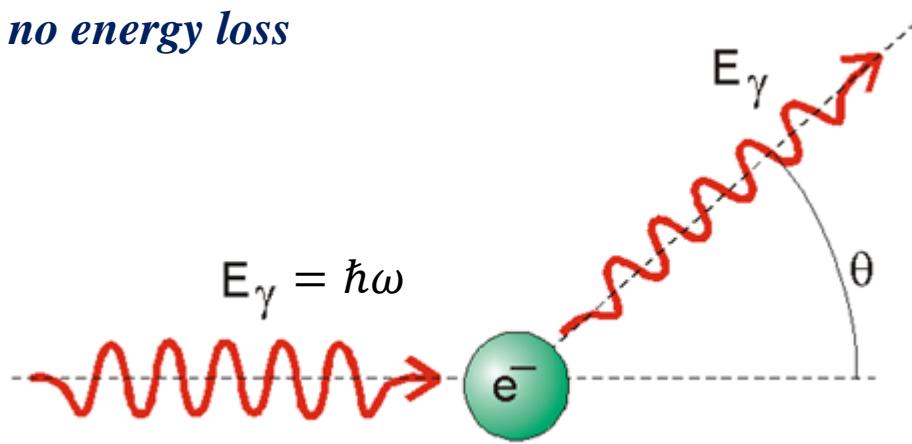
ATTENUATION

THOMSON SCATTERING

elastic scattering of el.-mag. waves at a free charges = electron, ...

independent of wave length λ

no energy loss



$$\sigma_{Th} = \frac{8\pi}{3} \cdot r_e^2$$
$$\cong \frac{2}{3} \text{ barn}$$

$$\sigma_{Th, atom} = Z \cdot \sigma_{Th}$$

deviates from experiment

$$r_e = \frac{e^2}{4\pi\epsilon_0 m_e c^2}$$

$$= \alpha \cdot \frac{\hbar c}{m_e c^2}$$

$$= 2.82 \cdot 10^{-15} \text{ m}$$

application: plasma diagnosis, polarization of CMB, ...

RAYLEIGH SCATTERING

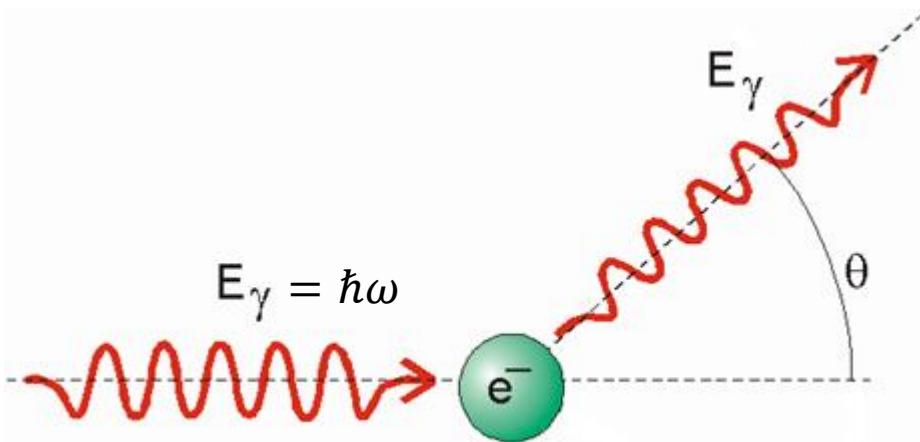
elastic scattering of el.-mag. waves at polarisable scattering centers = atoms, molecules

damped oscillation of „elastically“ bound electrons

eigen frequency ω_0 of bound system

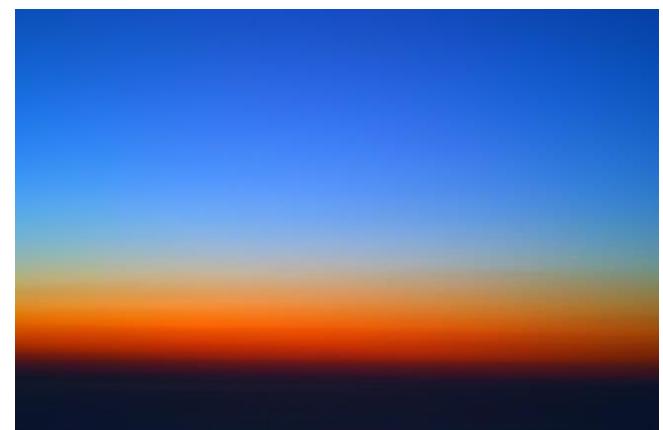
no energy loss

$$\sigma_R = \sigma_{Th} \cdot \frac{\omega^4}{(\omega^2 - \omega_0^2)^2} \cdot Z^2$$



application: combustion diagnosis, holidays, ...

$\omega \ll \omega_0$ makes the sky blue / sunset red

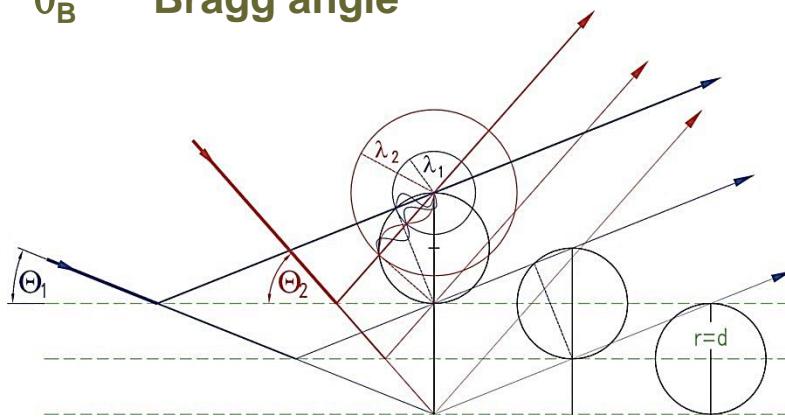


COHERENT SCATTERING - BRAGG'S LAW

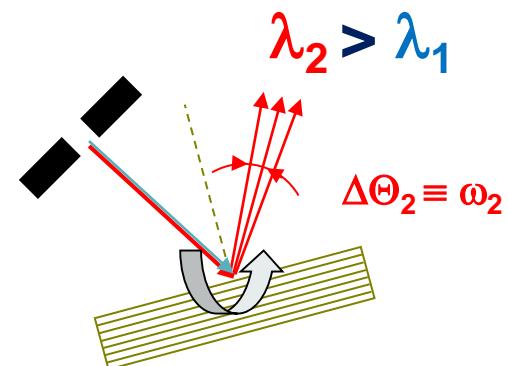
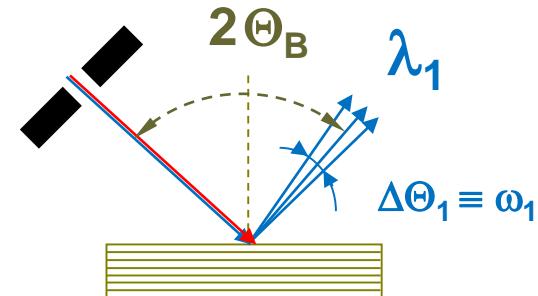
$$n\lambda = 2d \cdot \sin \theta_B$$

ω angular spread of „reflection“

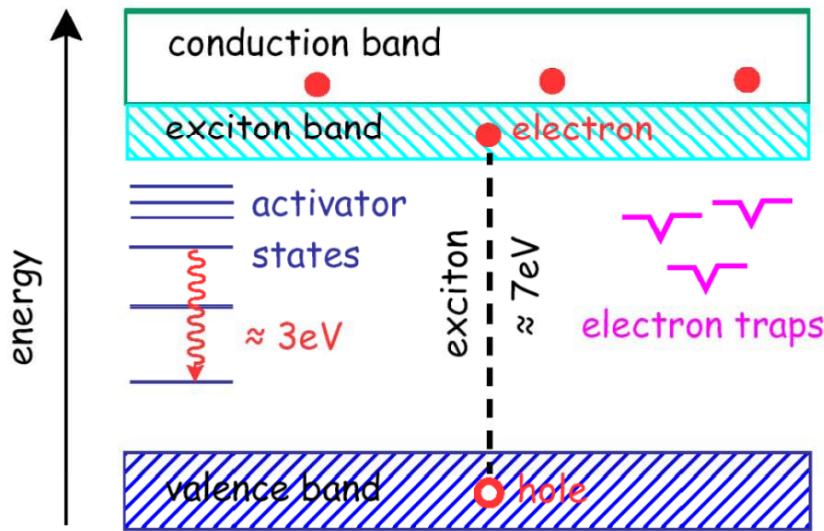
n order of diffraction
 λ wave length
 d spacing of diffracting planes
 θ_B Bragg angle



e.g. wave length of $\lambda = 60 \text{ pm}$
 electromagnetic wave $20 \text{ keV X-rays (dentist)}$
 matter wave $T = 0.2 \text{ eV}, v = 60 \text{ cm/s}$
 (epithermal neutrons)



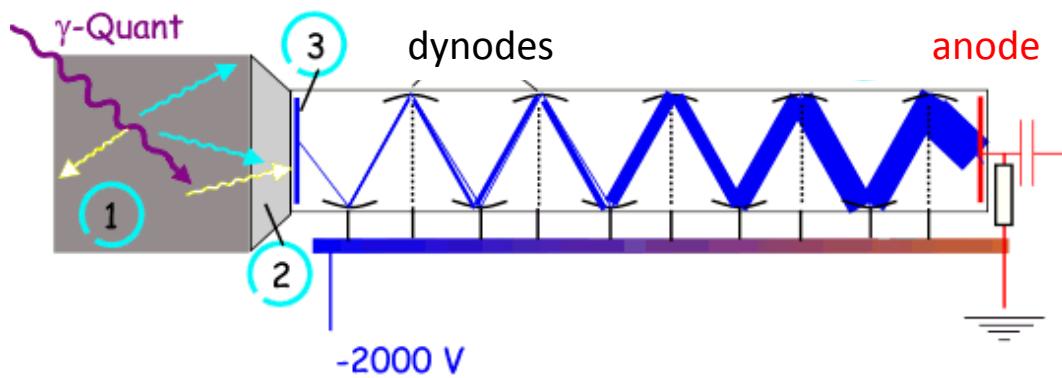
EXCITATION - SCINTILLATORS produce “LIGHT”



ionisation caused by
charged particles or light
excitation and delayed light emission
usually in the UV range

scintillators ①

inorganic NaI(Tl), CsI, BaF₂, ...
organic doped „plastics“



② light coupling

UV light is converted to charge
at a photo cathode ③ and
multiplied by a multi stage (dynodes)
photo „multiplier“

EXCITATION - RESPONSE of EYE CELLS

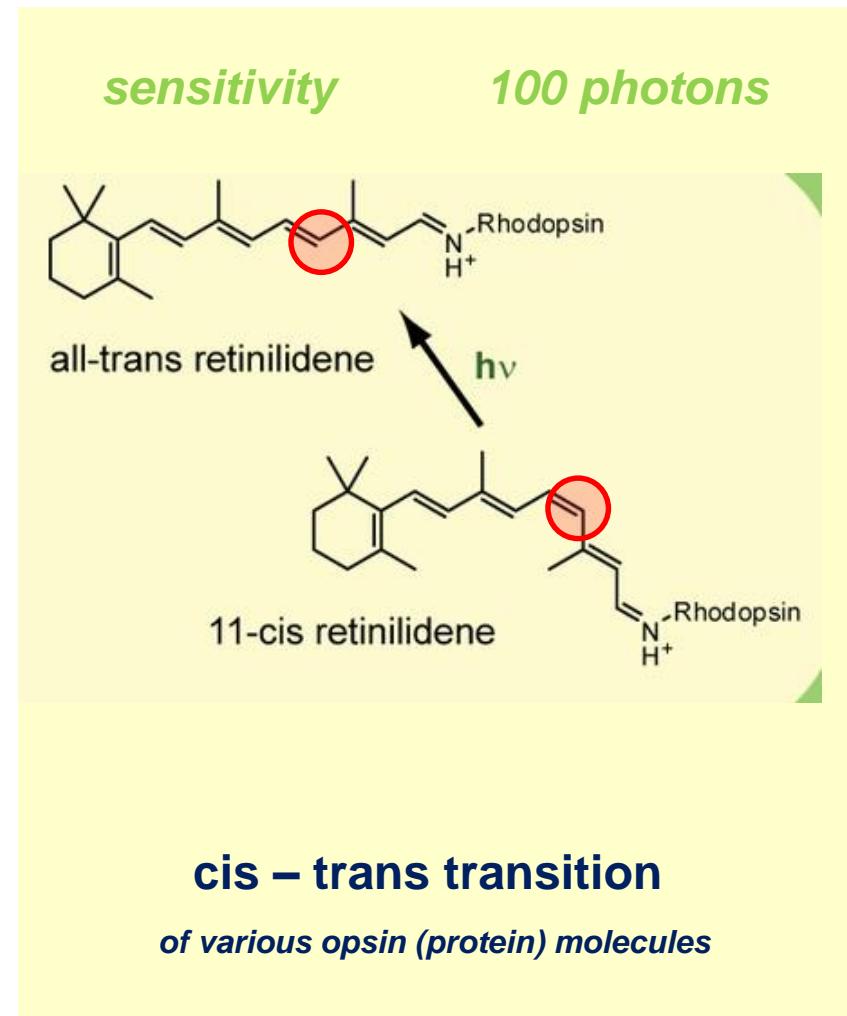
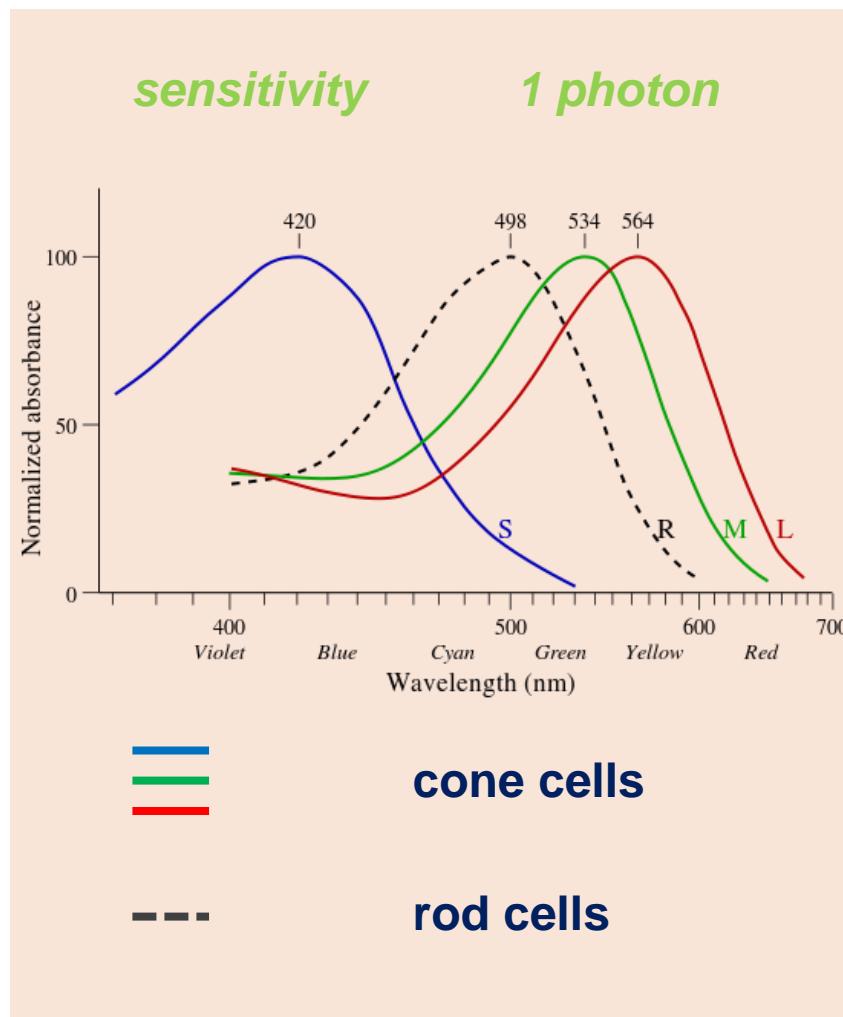


PHOTO EFFECT

requires particle nature of „light“ Einstein 1905

$$\sigma_{2K} = \sigma_{Th} \cdot 4\sqrt{2}\alpha^4 \cdot Z^{4-5} \cdot \varepsilon_\gamma^{-7/2}$$

1. photon disappears

photo electron $E_e = E_{\text{photon}} - E_B$

2. refilling of hole in electron shell by

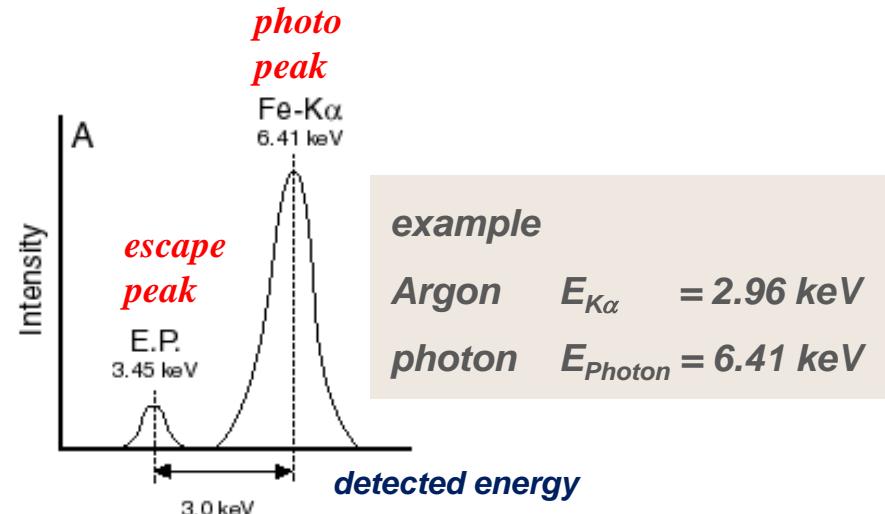
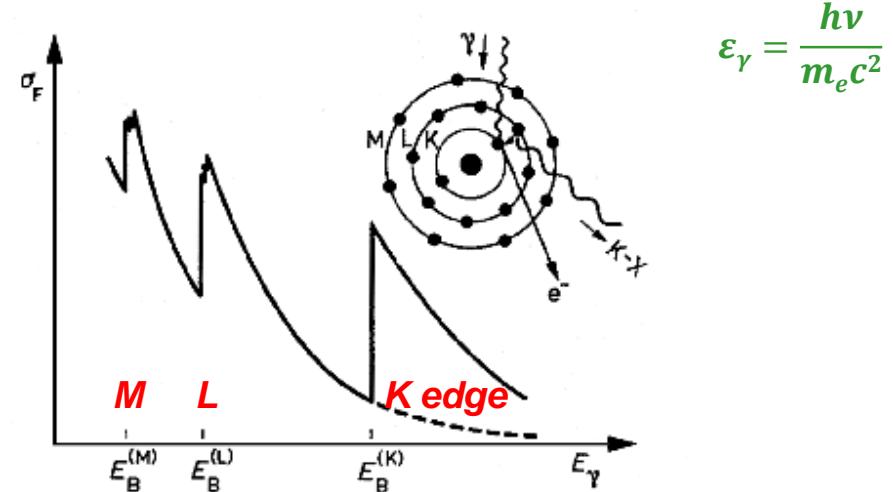
- a) emission of photon or
- b) Auger electron emission of loosely bound outer electron

$$E_{\text{Auger}} \approx E_B$$

detected energy E

photo peak $E = E_{\text{photon}}$
 $= E_e + E_B$

escape peak $E = E_{\text{photon}} - E_{K\alpha}$



COMPTON EFFECT

proof of particle nature of „light“ Compton 1922

billard with photons and „quasifree“ electrons

$$\sigma_C \approx \sigma_{Th} \cdot (1 - 2\epsilon\gamma + \dots) \cdot Z \quad \epsilon_\gamma \ll 1$$

$$\approx \sigma_{Th} \cdot \frac{3}{4} \cdot \left(\frac{1+2\ln\epsilon\gamma}{2\epsilon\gamma} + \dots \right) \cdot Z \quad \epsilon_\gamma \gg 1$$

complicated QED calculation Klein&Nishina 1929

photon does not disappear

recoil electron

$$E_e = E_{\text{photon}} - E'_{\text{photon}}$$

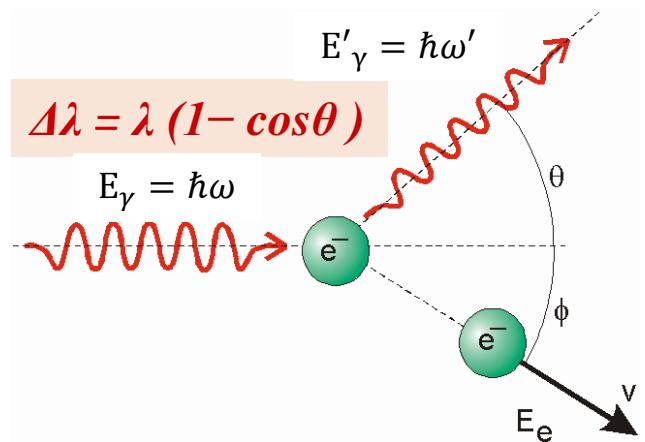
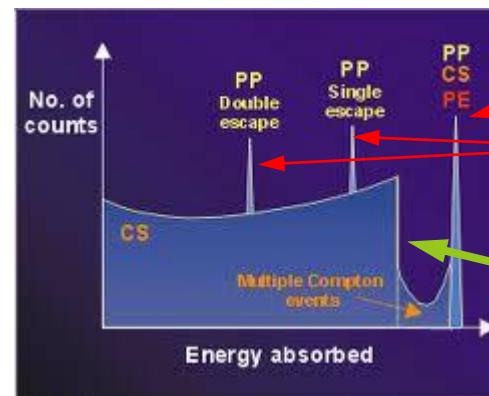


continuous spectrum

detected energy E

= absorbed energy E_e

we neglect E_B of the electron
and E_{recoil} of the nucleus
because usually $E_B, E_{\text{recoil}} \ll E_e$



$$\Delta\lambda = \lambda (1 - \cos\theta)$$

PAIR PRODUCTION

proof of mass-energy equivalence Blackett 1948

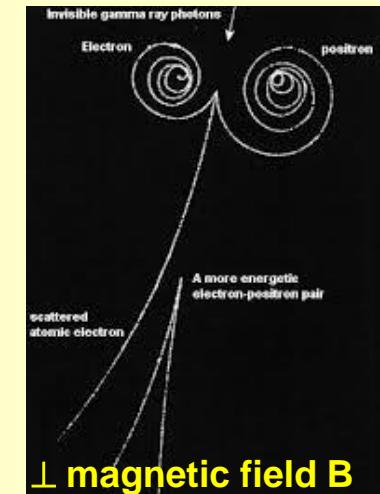
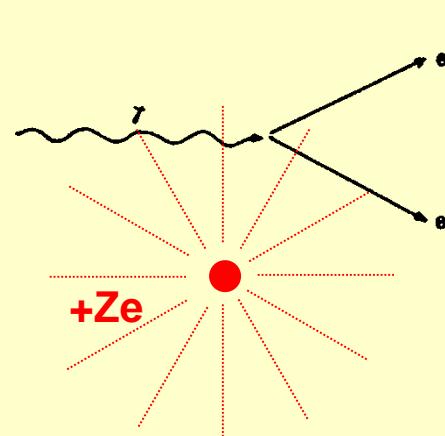
$$\sigma_{pair} \approx \sigma_{Th} \cdot Z^2 \cdot (\ln 2\epsilon\gamma + \dots) \quad \epsilon\gamma \gg 1$$

conversion of energy into matter

$$E_{\text{photon}} = h\nu > 2 m_{\text{electron,muon,pion, ...}}$$

*a recoil partner (e.g. a nucleus) is needed
to fulfil energy and momentum conservation*

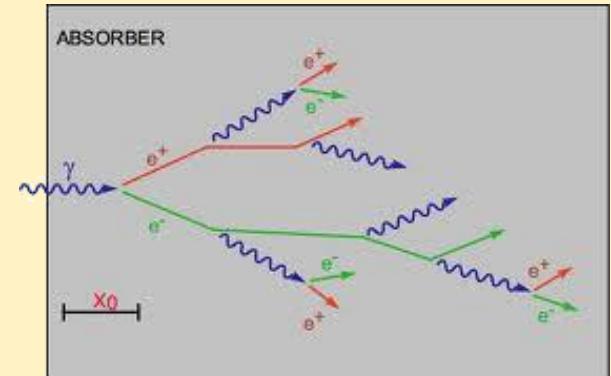
$$\begin{aligned} e^+e^- \text{ threshold: } m_{\text{recoil}} &= \infty \quad h\nu = 2 m_e c^2 \\ &= m_e \quad = 4 m_e c^2 \end{aligned}$$



el.-mag. shower

$e^+ e^- \gamma$ - cascade
pair production and bremsstrahlung alternate
shower may start with photon or electron

radiation length X_0
characteristic material dependent constant
depth, where about 2/3 (1/e) of the incident energy is converted



BREMSSTRAHLUNG

accelerated charged particles radiate Hertz 1886

electromagnetic waves

bending force by Coulomb potential

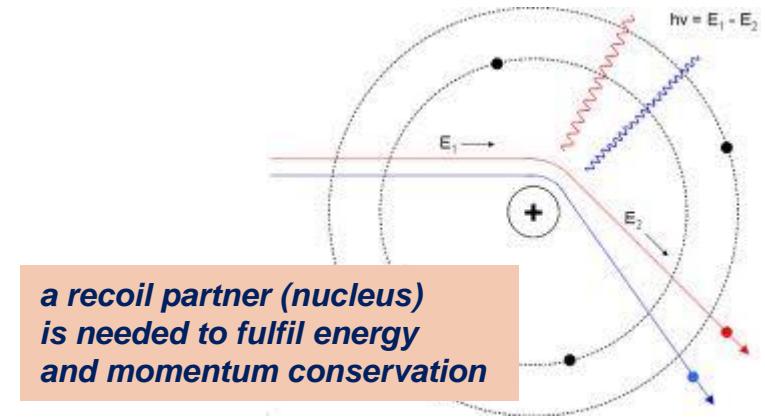
force \Leftrightarrow acceleration

$$F_{\text{Coulomb}} = \frac{1}{4\pi\epsilon_0} \cdot \frac{Q_{\text{particle}} \cdot Q_{\text{nucleus}}}{r^2}$$
$$= m \cdot \ddot{r}$$

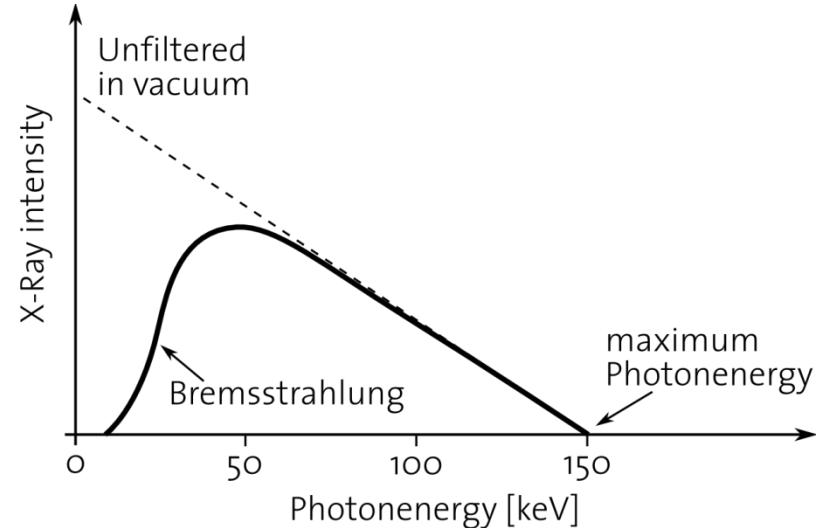
any distance r

\Rightarrow continuous spectrum

$$\sigma_b \approx \sigma_{Th} \cdot Z^2 \cdot [\text{energy dependent}]$$



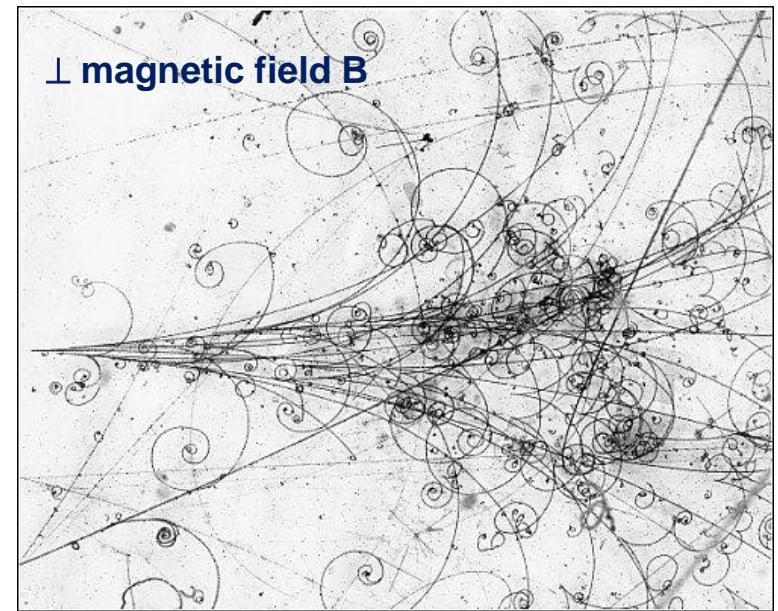
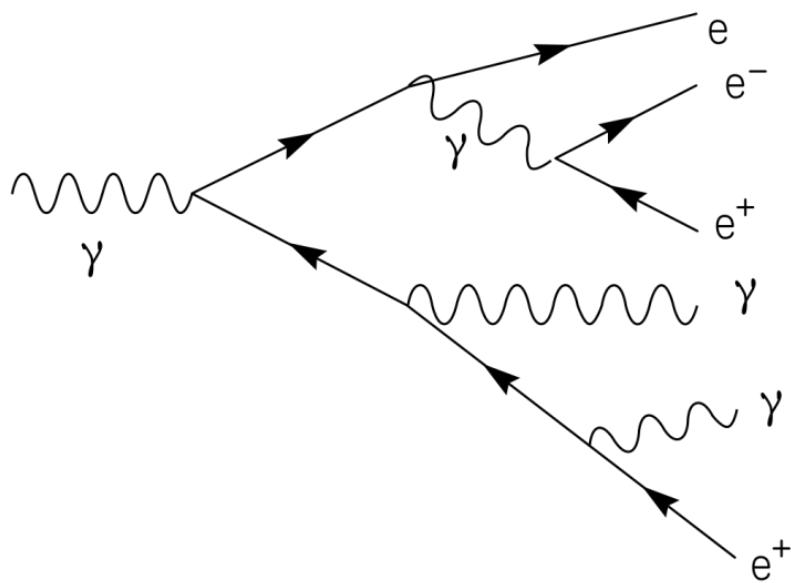
a recoil partner (nucleus)
is needed to fulfil energy
and momentum conservation



EL.-MAG. SHOWER

alternating pair production & bremsstrahlung

initial particle of minor importance for large energies

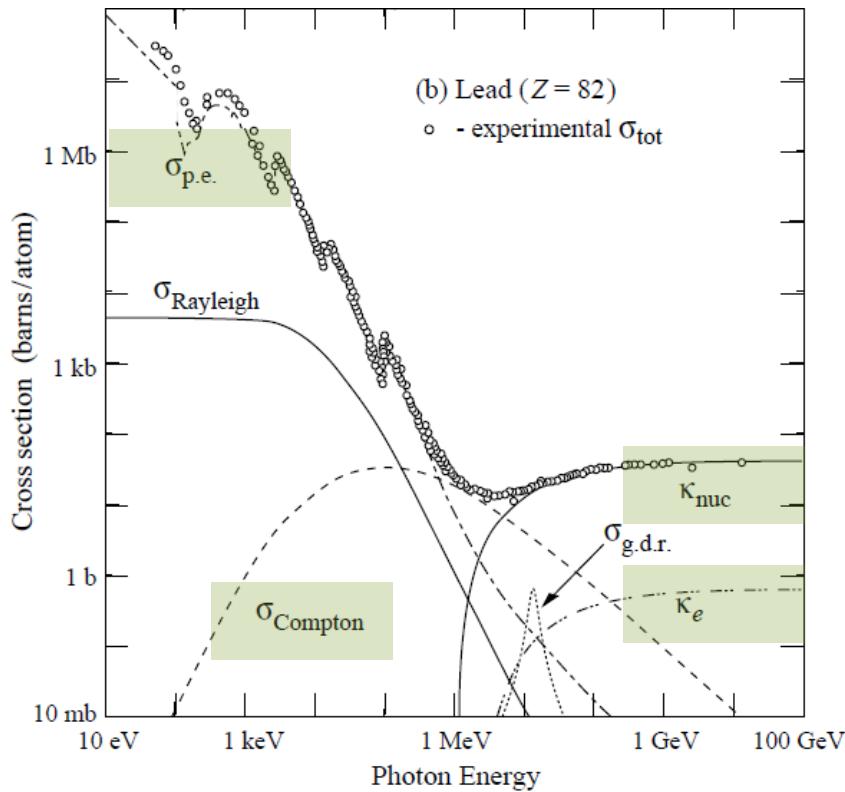


radiation length X_0

characteristic quantity of absorber

$$E_\gamma = E_{\text{initial}} \cdot e^{-(x/X_0)}$$

CROSS SECTIONS SUMMARY



from
[C. Patrignani et al.](#)
[\(Particle Data Group\)](#),
 Chin. Phys. C, **40**, 100001 (2016).

$\sigma_{\text{p.e.}}$ = Atomic photoelectric effect (electron ejection, photon absorption)

σ_{Rayleigh} = Rayleigh (coherent) scattering—atom neither ionized nor excited

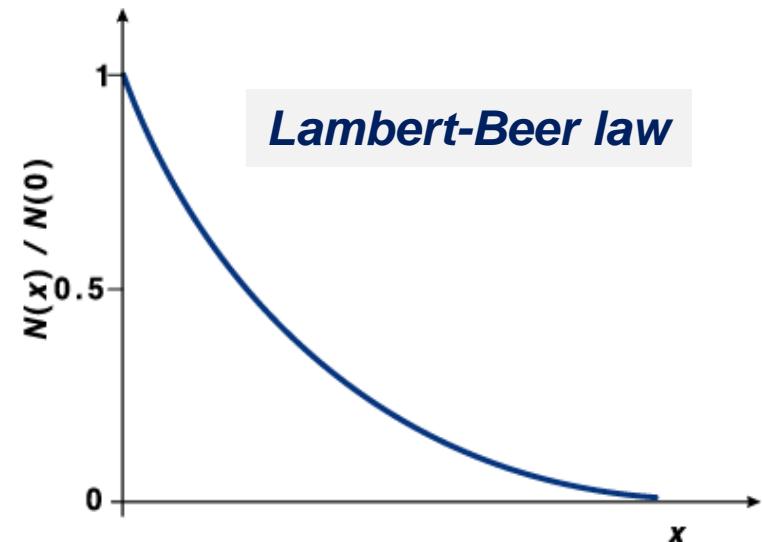
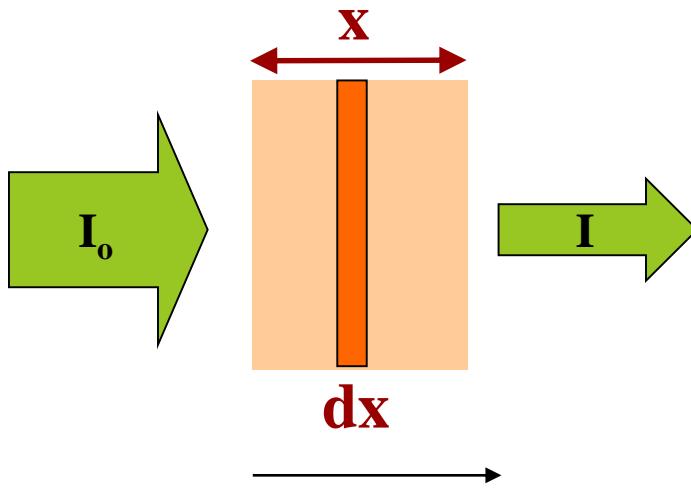
σ_{Compton} = Incoherent scattering (Compton scattering off an electron)

κ_{nuc} = Pair production, nuclear field

κ_e = Pair production, electron field

$\sigma_{\text{g.d.r.}}$ = Photonuclear interactions,

ATTENUATION



$$I(x) = I_0 e^{-\mu(hv)x} \quad \text{intensity after layer thickness } x$$

$$\frac{I_T(x)}{I_0} = e^{-\mu(hv)x} \quad \text{fraction of transmission}$$

$$\frac{I_A(x)}{I_0} = 1 - e^{-\mu(hv)x} \quad \text{fraction of absorption}$$

sum of linear attenuation coeff.

$$\mu(hv) = \sum_i \mu_i(hv)$$

$$\mu_i = \rho \cdot \frac{N_A}{A} \cdot \sigma_i(hv)$$

INTERACTION OF

MASSIVE PARTICLES

CHARGED PARTICLES : ENERGY LOSS BY IONIZATION

HEAVY CHARGED PARTICLES

LIGHT CHARGED PARTICLES

CHARGED PARTICLES : ENERGY LOSS BY RADIATION

NEUTRONS

CHARGED PARTICLES

interaction happens by collisions of particles type 1 and 2

1. $M_{\text{particle 1}} \gg M_{\text{particle 2}}$

before



after collision



2. $M_{\text{particle 1}} = M_{\text{particle 2}}$



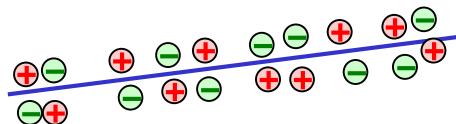
CHARGED PARTICLES - ENERGY LOSS by IONIZATION

collisions create electron- ion pairs

1. **heavy**

$$M_{\text{particle}} \gg M_{\text{electron}}$$

e.g. protons, deuterons, ...



strongly ionising

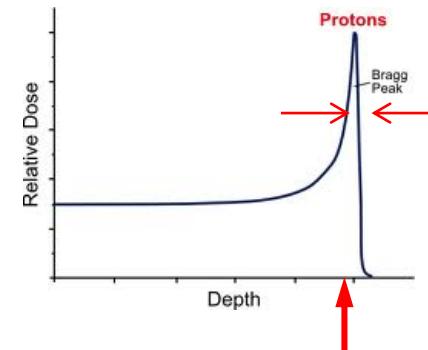
2. **light**

$$M_{\text{particle}} = M_{\text{electron}}$$

electrons or positrons



weakly ionising

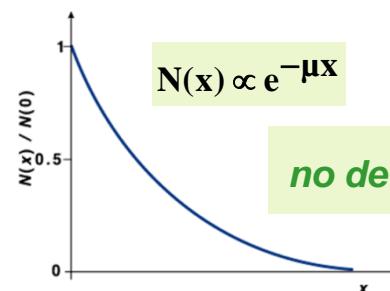


Bragg peak

$$\frac{\Delta R}{R} = 1 - 3\%$$

for all elements

well defined range R!



no defined range R!

exponential attenuation with depth x

μ : material dependent attenuation coefficient

INTERACTION OF

HEAVY CHARGED PARTICLES

WITH MATTER

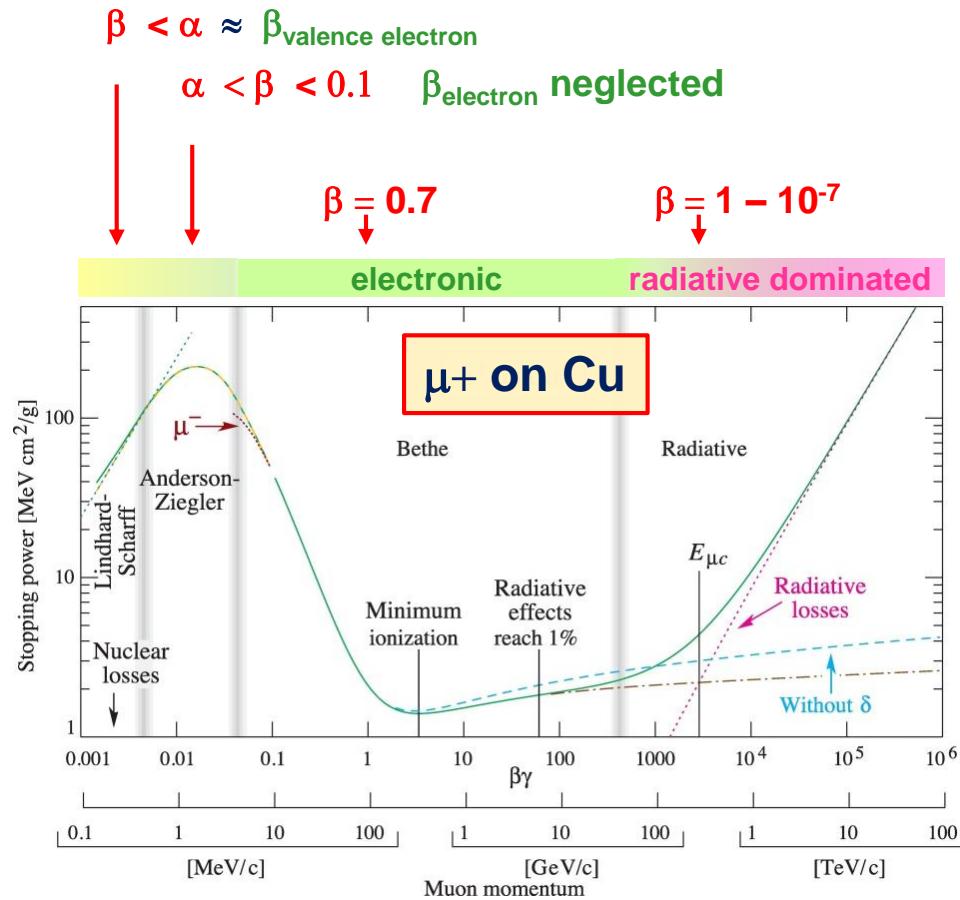
HEAVY CHARGED PARTICLES - STOPPING POWER I

heavy particles μ, π, K, p, d, \dots

stopping power

$$S = \left(-\frac{dE}{dx} \right) \cdot \frac{1}{\rho}$$

[MeV · cm²/g]



from
[C. Patrignani et al.](#)
[\(Particle Data Group\)](#),
 Chin. Phys. C, **40**, 100001 (2016).

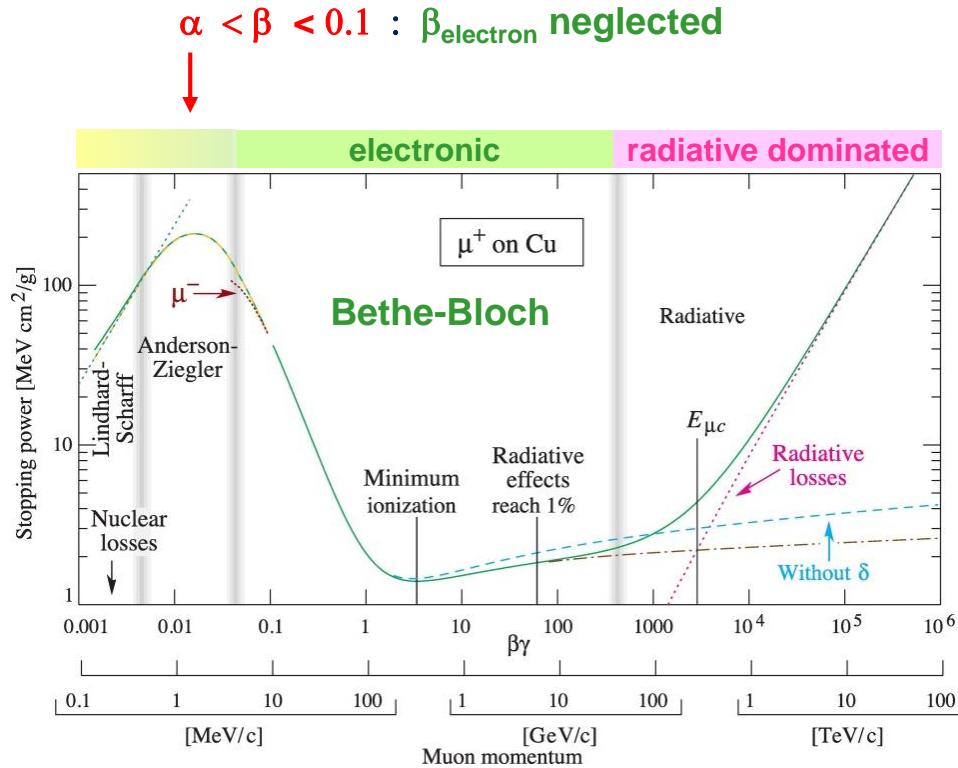
HEAVY CHARGED PARTICLES - STOPPING POWER II

Bethe-Bloch range

stopping power

$$S = \left(-\frac{dE}{dx} \right) \cdot \frac{1}{\rho}$$

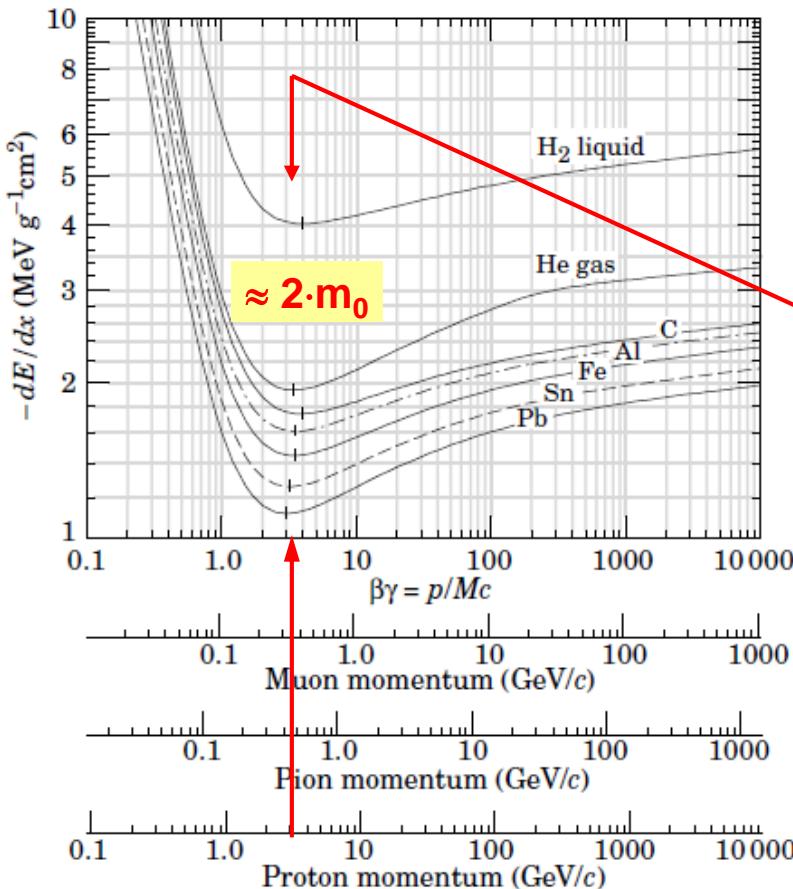
[MeV · cm²/g]



C_0 shell corrections
for low energies

$$S = 4\pi N_A r_e^2 m_e c^2 \frac{Z_{\text{target}}}{A} \cdot \frac{z_{\text{projectile}}^2}{\beta^2} \cdot \left[\frac{1}{2} \ln \left(\frac{2m_e c^2 \beta^2 \gamma^2 T_e^{\max}}{I^2} \right) - \beta^2 - \frac{C_0(\beta^2)}{Z_{\text{target}}} - \frac{\delta(\beta\gamma)}{2} \right]$$

HEAVY CHARGED PARTICLES - STOPPING POWER III



Bethe-Bloch range

$$\left(\frac{\Delta E}{\Delta X} \right)_{\text{collision}} \propto \frac{1}{v^2} \dots$$

MIPs = *minimum ionising particles*

$$\left(-\frac{dE}{dx} \right)_{\min} = 4 \text{ MeVg}^{-1}\text{cm}^2 \quad \text{for } A = 1$$

$$\left(-\frac{dE}{dx} \right)_{\min} = 1 - 2 \text{ MeVg}^{-1}\text{cm}^2 \quad \text{for } A > 1$$

$$\text{at } \beta\gamma = 3 - 3.5 \text{ or } T_{\min} = 2.2 - 2.6 m_0 c^2$$

Figure 30.2: Mean energy loss rate in liquid (bubble chamber) hydrogen, gaseous helium, carbon, aluminum, iron, tin, and lead. Radiative effects, relevant for muons and pions, are not included. These become significant for muons in iron for $\beta\gamma \gtrsim 1000$, and at lower momenta for muons in higher- Z absorbers. See Fig. 30.23.

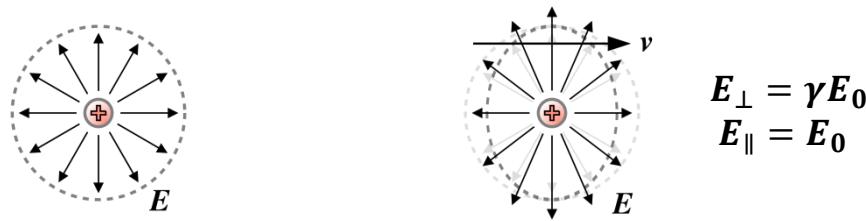
from
[C. Patrignani et al.](#)
[\(Particle Data Group\)](#),
 Chin. Phys. C, **40**, 100001 (2016).

HEAVY CHARGED PARTICLES - STOPPING POWER IV

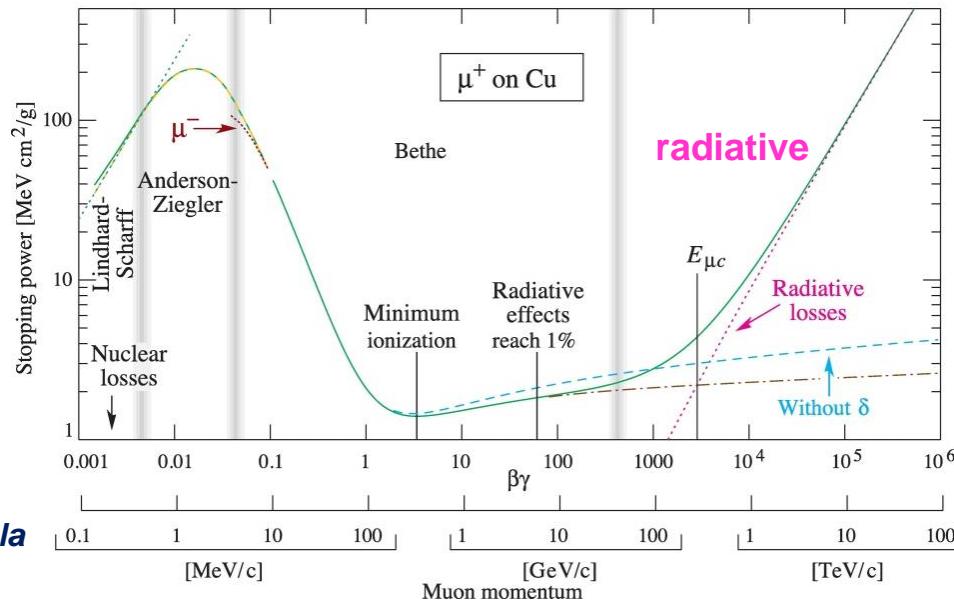
„LHC“ range

radiative losses

- bremsstrahlung
- pair production e^+e^-
- photonuclear



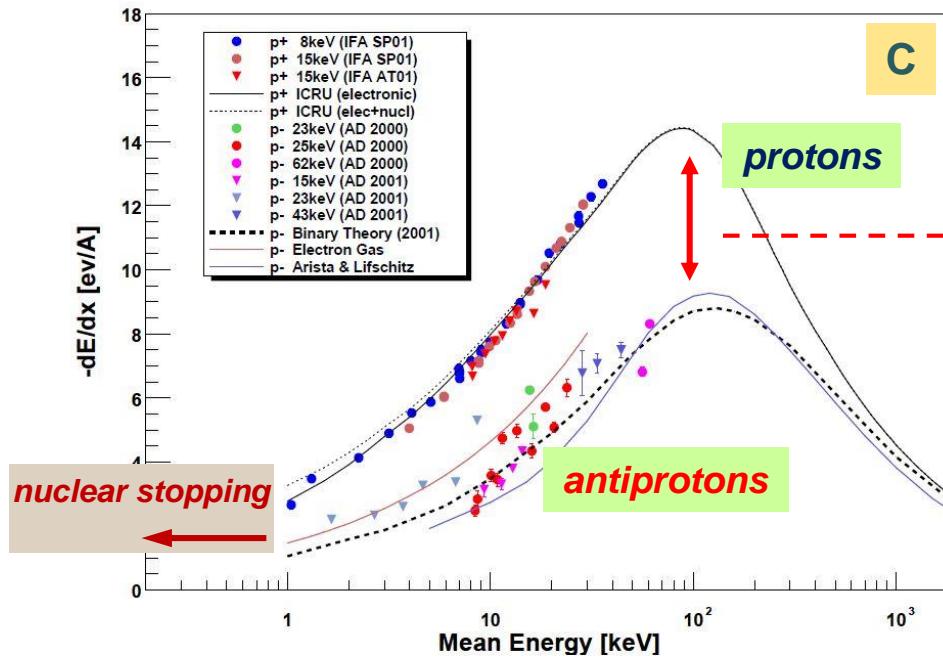
not covered by this formula



δ density effect
polarization diminishes relativistic rise

$$S = 4\pi N_A r_e^2 m_e c^2 \frac{Z_{target}}{A} \cdot \frac{Z_{projectile}^2}{\beta^2} \cdot \left[\frac{1}{2} \ln \left(\frac{2m_e c^2 \beta^2 \gamma^2 T_e^{max}}{I^2} \right) - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$$

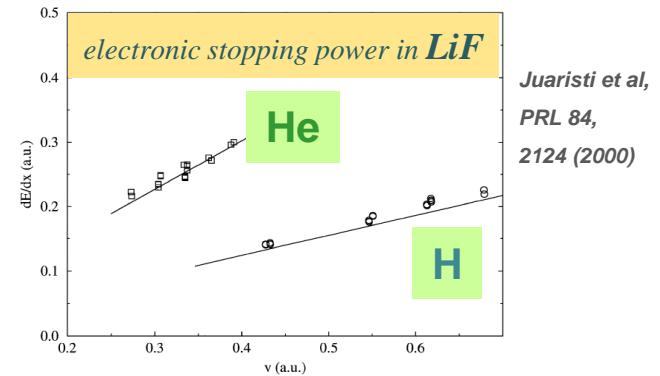
HEAVY CHARGED PARTICLES - BARKAS EFFECT



lowest energies - friction range

like friction

$$-\left(\frac{\Delta E}{\Delta x}\right) \propto v$$



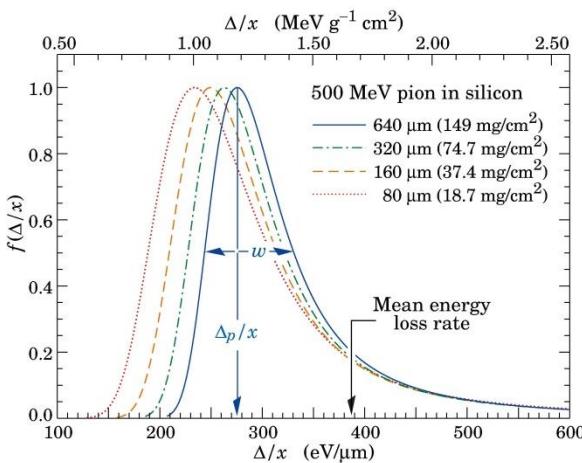
sensitive to sign of charge

$$S = 4\pi N_A r_e^2 m_e c^2 \frac{Z_{target}}{A} \cdot \frac{Z_{projectile}^2}{\beta^2} \cdot \{ [...] + L_1(\beta, Z_{target}) \cdot z_{projectile} \}$$

frictional cooling (e-cooler, muon collider), window design, ...

HEAVY CHARGED PARTICLES - STRAGGLING

energy (loss) straggling Δ



Landau-Vavilov distribution

asymmetric energy straggling towards higher Δ

thick layers → many collisions → skewness decreases

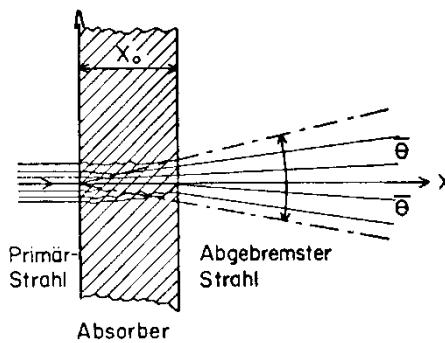
Δ_p/x *most probable energy loss (here normalized to unity)*

Δ/x *energy loss per layer thickness*

$$\overline{\Delta}^2 \propto \frac{Z}{A} \cdot \rho \cdot d \cdot \frac{1}{\beta^2} \quad \text{for „thin“ layers}$$

from
[C. Patrignani et al.](#)
[\(Particle Data Group\)](#),
 Chin. Phys. C, **40**, 100001 (2016).

angular straggling $\bar{\Theta}$



$$\overline{\Theta} = \frac{13.6 \text{ MeV}}{\beta c p} \sqrt{x / X_0} (1 + \dots) \propto z \cdot Z / p^2$$

many collisions → Gaussian angular distribution

$$\begin{aligned} X_0 / \text{g cm}^{-2} &= 63 \text{ (126)} \quad \text{H}_2 (\text{D}_2) \quad \text{radiation length} \\ &= 108 \quad \text{Si} \\ &= 13.8 \quad \text{Fe} \end{aligned}$$

$x / \text{g cm}^{-2}$ *effective thickness of layer ($x = d \cdot \rho$)*

- *acceptance of experimental setup (storage rings etc.)*
- *position resolution of tracking devices*

HEAVY CHARGED PARTICLES - RANGE I

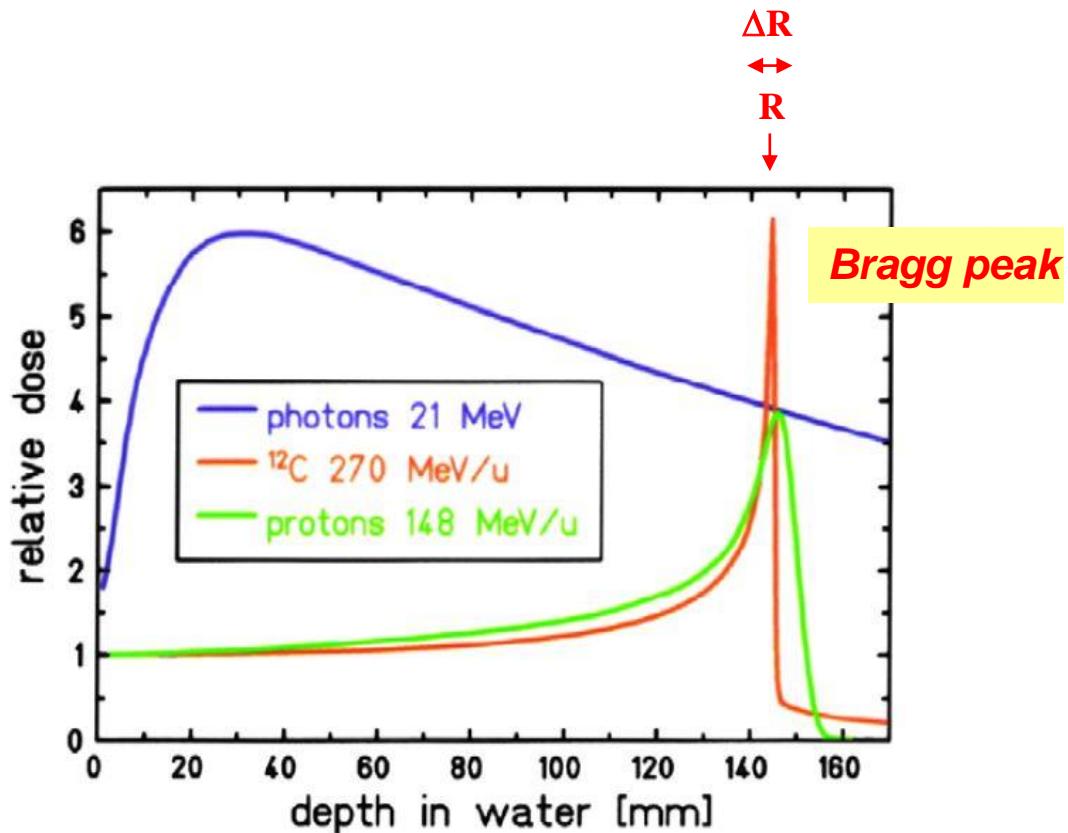


Fig. 1. Depth dose distribution for photons and monoenergetic Bragg curves for carbon ions and protons (Courtesy of G. Kraft, GSI Darmstadt, Germany).

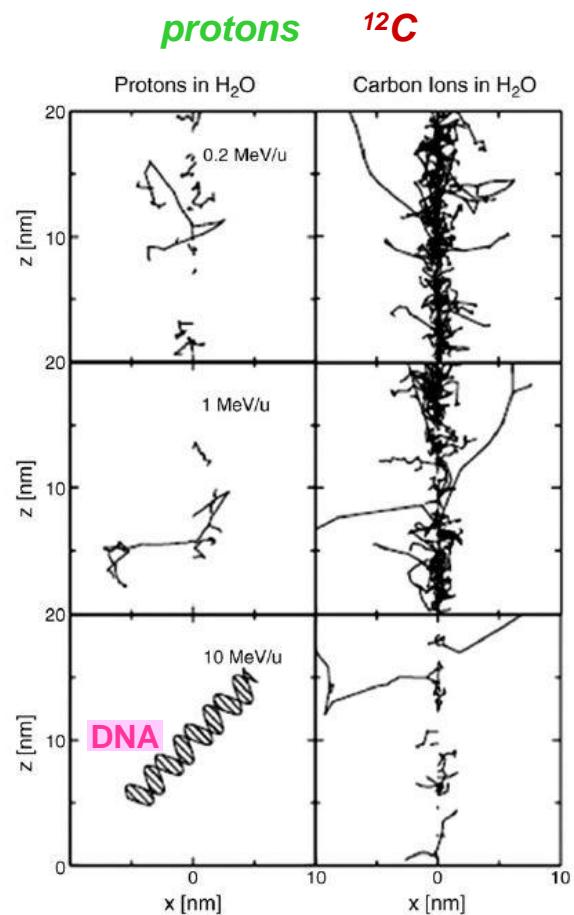


Fig. 4. Proton and carbon ion tracks are compared microscopically to an illustration of a DNA molecule before, in and behind the Bragg maximum, for the same energy [41].

HEAVY CHARGED PARTICLES - RANGE II

mean range
depends on particle mass

$$R = \int_0^{\infty} dE / (dE/dx) \quad [\text{cm}]$$

$$T_{\text{kin}}$$

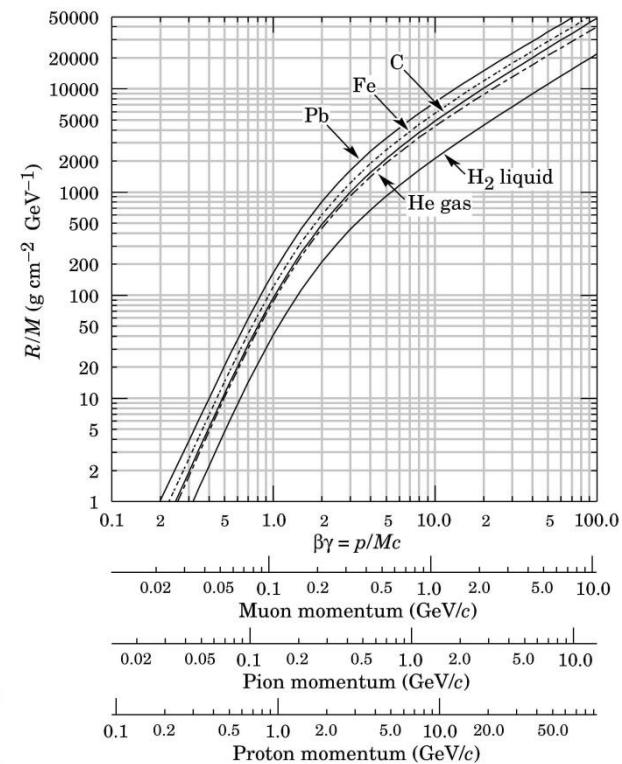
range – straggling

$$\Delta R$$

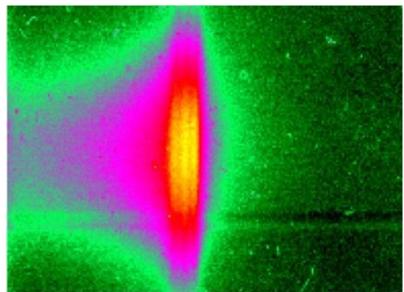
longitudinal
transversal

$$\Delta R/R \approx 1\% - 3\% \quad \text{for all elements}$$

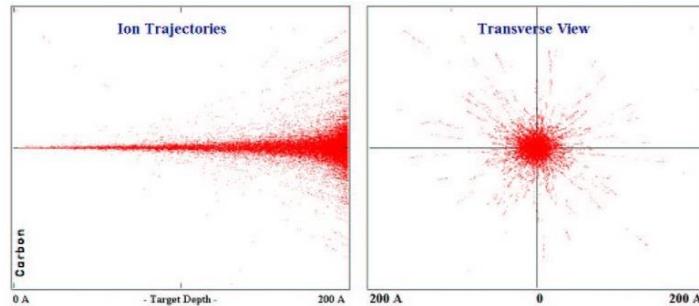
$$\approx 2\% - 6\%$$



47 MeV antiprotons
radiochromic film response



20 keV protons on carbon
(Monte-Carlo simulation SRIM)



N. Bassler et al.
Radiotherapy and Oncology 86 (2008) 14–19

A.Csete / PhD thesis, Aarhus, 2002

R/M(E/M)

range concept useful for

- $R < \lambda_{\text{had}}$
- **radiation losses small**

INTERACTION OF

LIGHT CHARGED PARTICLES

WITH MATTER

LIGHT CHARGED PARTICLES - STOPPING POWER

ionisation dominated energy range

heavy particles

path lengths $S \approx \text{range } R$

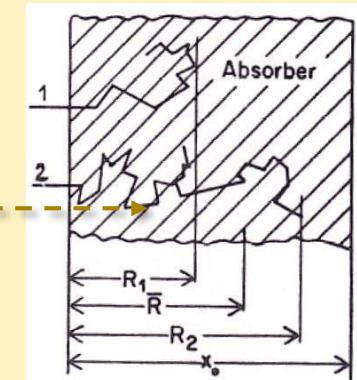
electrons/positrons

" " $S \approx 2 \cdot R$

strong deflection

dE/dx similar to Bethe-Bloch formula

additional terms
 - identical particles (e^-)
 - spin dependence



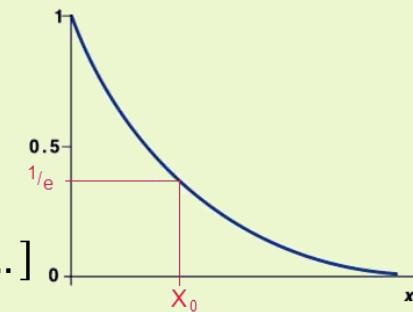
radiation dominated energy range

energy loss by bremsstrahlung $-\frac{dE_{kin}}{dx} \propto Z_{target}^2 \cdot E_{kin} \cdot [\dots]$

$$\Rightarrow E_{kin} = E_{0,kin} \cdot e^{-(x/x_0)}$$

radiation length $X_0 [\text{g}\cdot\text{cm}^{-2}]$

$$\frac{1}{X_0} = 4 \alpha \cdot r_e^2 \cdot \frac{N_A}{A} \cdot Z_{target}^2 \cdot [\dots]$$



after depth $d = X_0 / \rho$ ([cm]) all but $1/e$ of the energy of the particle is lost by bremsstrahlung

LIGHT CHARGED PARTICLES - RANGE

ionisation dominated energy range

electron range (semiempirical formulae)

$$R = 0,52 \frac{E^{(\text{MeV})}}{E_e} - 0,09 \quad (\text{g cm}^{-2}) \quad 0,5 < E_e < 3 \text{ MeV}$$

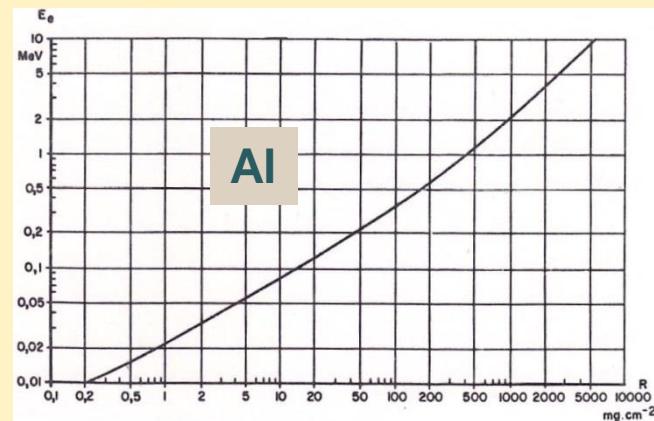
$$R = 0,412 E^n \quad (\text{g cm}^{-2}) \quad 0,01 < E_e < 3 \text{ MeV}$$

mit $n = 1,265 - 0,0954 \ln E_e$

$$R = 0,53 \frac{E^{(\text{MeV})}}{E_e} - 0,106 \quad (\text{g cm}^{-2}) \quad 1 < E_e < 20 \text{ MeV}$$

$$-\frac{dE}{dx} = \frac{2 \pi e^4}{m_e c^2} N^a Z \left(\ln \frac{E_e}{I} + 0,15 \right) \quad E_e \ll m_e c^2$$

$$-\frac{dE}{dx} = \frac{2 \pi e^4}{m_e c^2} N^a Z \left(\ln \frac{E_e^3}{2m_e^2 I^2} + \frac{1}{8} \right) \quad E_e \gg m_e c^2$$



radiation dominated energy range

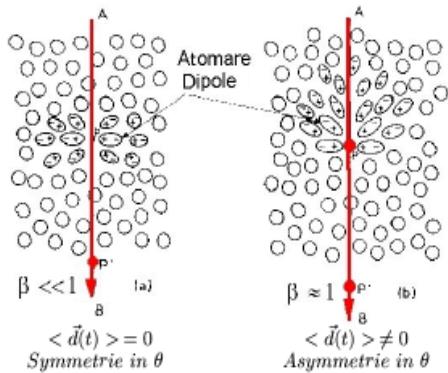
radiation length X_0 [g·cm⁻²]

D ₂	126	mylar	40
H ₂	63	air	37
Al	24	water	36
Ar	20	rock standard	27
Cu	13	Csl	8.4
Pb	6	PbWO ₄	7.4

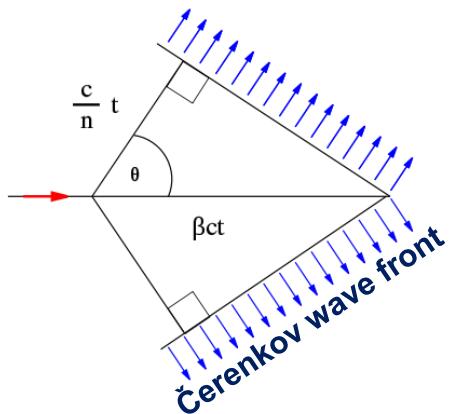
CHARGED PARTICLES - ENERGY LOSS BY RADIATION I

Čerenkov radiation if $v_{\text{particle}} > c_{\text{in medium}}$

the charge polarizes the medium

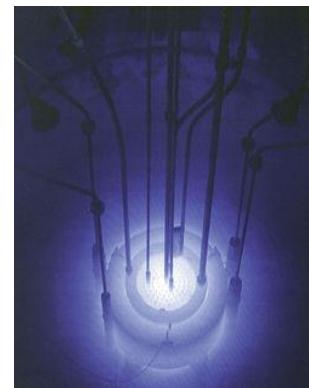


emission under specific angle $\Theta_{\text{č}}$



Čerenkov 1930s

„light“ blue!



electrons „radiate“
in the water above
the core of
a nuclear power plant

$$\left(\frac{\Delta E}{\Delta x} \right)_{\text{Čerenkov}} \ll \left(\frac{\Delta E}{\Delta x} \right)_{\text{collision}}$$

$$\cos \Theta_{\text{č}} = 1 / \beta \cdot n$$

$n = \text{index of refraction}$

(small) dispersion !

$\Theta_{\text{č}}$ measures the velocity of the particle

acoustics analogue: Mach's cone for supersonic source

CHARGED PARTICLES - ENERGY LOSS BY RADIATION II

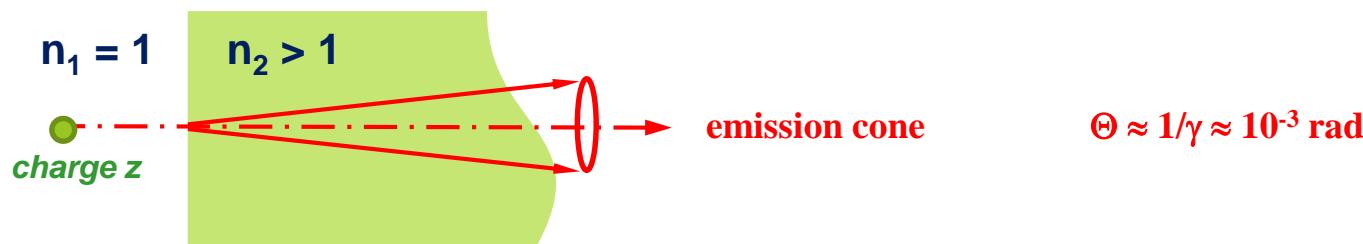
Transition radiation for ultrarelativistic particles ($\gamma \gg 1$)

Ginzburg & Frank 1946

Readjustment of the el.-mag. fields (E,H) at the boundary of 2 media

with different dielectric properties (ϵ)

leads as collective response of the material to emission of el.-mag. radiation (X-rays)



$$\Theta \approx 1/\gamma \approx 10^{-3} \text{ rad}$$

radiated intensity $I = \alpha \cdot z^2 \cdot \gamma \cdot \hbar \omega_p / 3$

photon yield $n_{\text{photon}} \propto \alpha \cdot z^2 \cdot (\ln \gamma)^2 \approx z^2 \cdot 0.5\%$

plasma frequency $\omega_p^2 = \frac{e^2}{\epsilon_0} \cdot \frac{n_e}{m_e}$ air: $\hbar \omega_p = 0.7 \text{ eV}$

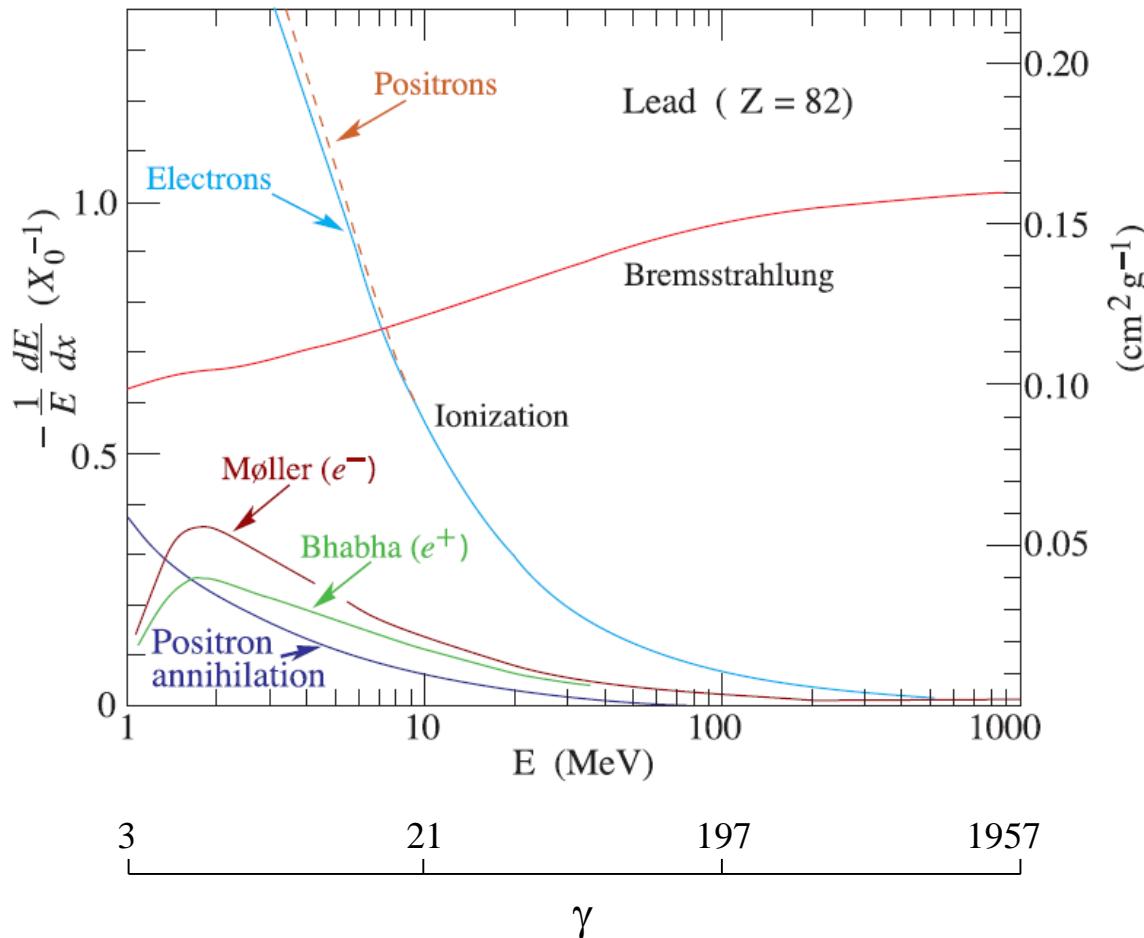
mylar: $\hbar \omega_p = 20 \text{ eV}$

formation length $d = \frac{\gamma}{\sqrt{2}} \cdot \frac{\hbar c}{\hbar \omega_p}$ mylar: $d = 14 \mu\text{m}$ ($\gamma = 1000$)

typical: soft X-rays of 2 - 40 keV for $\gamma \approx 1000$

application: plasma frequencies of materials, particle separation (π/p), ...

LIGHT CHARGED PARTICLES - RELATIVE ENERGY LOSS



from
[C. Patrignani et al.](#)
[\(Particle Data Group\)](#),
 Chin. Phys. C, **40**, 100001 (2016).

Fractional energy loss per radiation length in lead as a function of electron or positron energy.

INTERACTION OF

MASSIVE NEUTRAL PARTICLES

WITH MATTER

NEUTRONS I

collisions create recoil particles

maximum energy transfer for $M_{neutral} = M_{recoil}$

central collision energy is transferred completely

non central all energies according to scattering angle

average energy transfer 50%



detection by recoil protons (from hydrogen)

$$M_{proton} \approx M_{neutron}$$

i.e. good shielding is water - H_2O

concrete - 15% water

paraffin - $(CH)_n$

...

cloud chamber picture



neutron

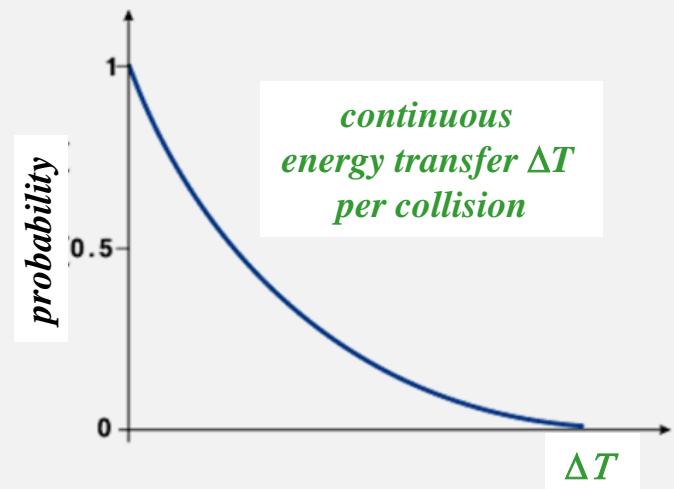
NEUTRONS II

slowing down of neutrons in elastic collisions

$$\left(\frac{A-1}{A+1}\right)^2 \cdot T_n \leq T_{n'} \leq T_n$$

T_n initial kinetic energy

$T_{n'}$ kinetic energy after collision



neutrons – no defined range

make degrader thick enough to thermalize neutrons, i.e. $T_n \approx \frac{1}{40} \text{ eV}$

subsequent capture or decay

don't forget absorber for reaction and decay products (mostly γ)

SYNOPSIS:

***BASIC INTERACTIONS ARE
BASIS FOR DETECTOR DESIGN***